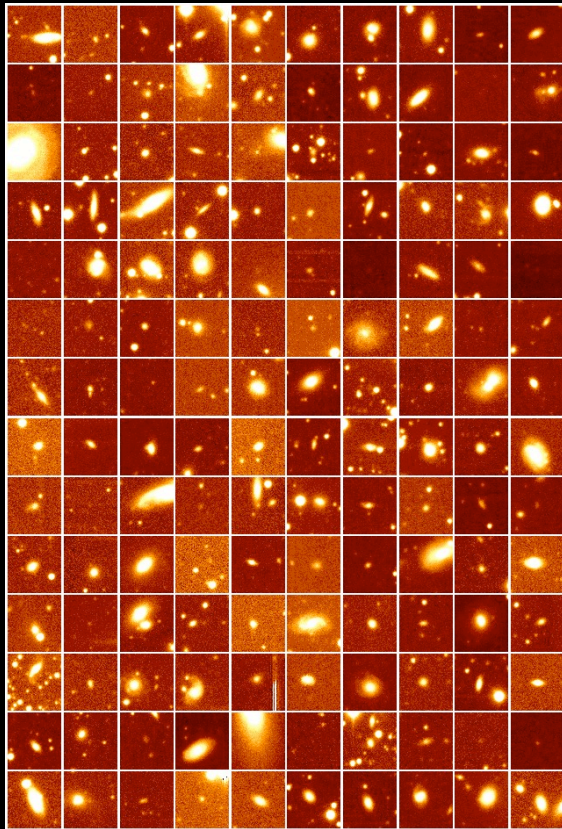
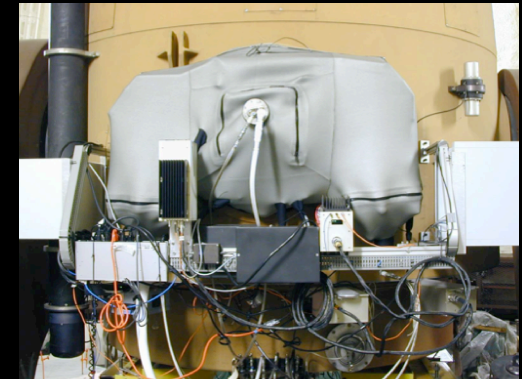
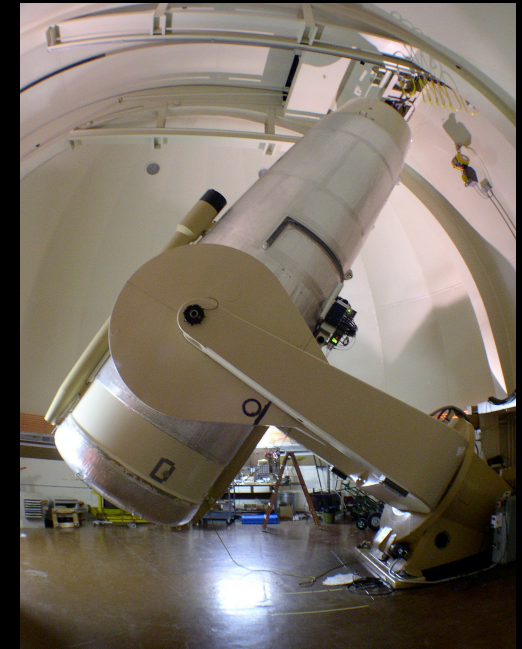
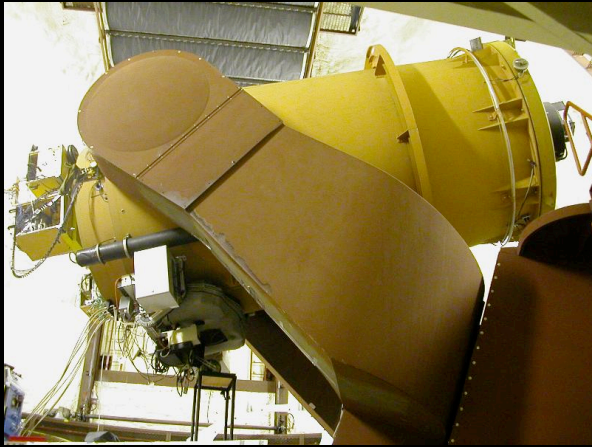


# Nearby Supernova Factory



## IPNL

C. Buton  
N. Chotard  
Y. Copin  
E. Gangler  
G. Smadja  
R. Pereira

## LPNHE

P. Antilogus  
S. Bailey  
S. Bongard  
R. Pain  
C. Wu

## Yale

C. Baltay  
D. Rabinowitz  
R. Scalzo

## LBL

G. Aldering  
M. Childress  
H. Fakhouri  
S. Loken  
S. Perlmutter  
P. Nugent  
R. Thomas  
K. Runge  
J. Zylberberg

## CRAL

E. Pecontal

## CPPM

C. Tao  
D. Fouchez

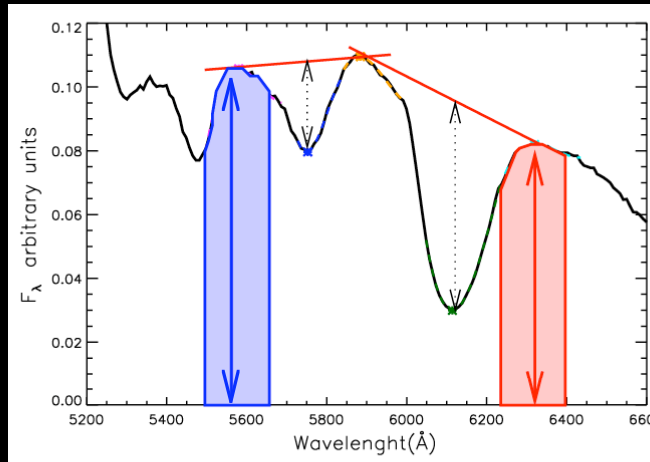
# A New Paradigm

- Search at low redshift just like at high redshift
  - boost discovery rate
  - wide-field, impartial to presence of an associated galaxy
  - find better analogs to high-redshift environment
- Fully integrate photometry and spectroscopy
  - *for photometry:*
    - Eliminate K-corrections and S-corrections
    - Tackle dust extinction using more colors
    - Eliminate atmospheric dispersion degeneracies
  - *for spectroscopy:*
    - Spectroscopy for more SNe
    - Time series of spectra for each SN
    - Properly subtract host galaxy light
    - Eliminate effects of slit losses and atmospheric dispersion

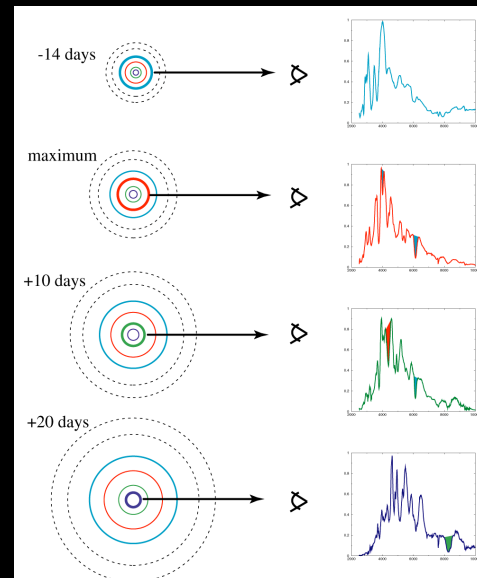


# Getting to the Physics with Spectroscopy

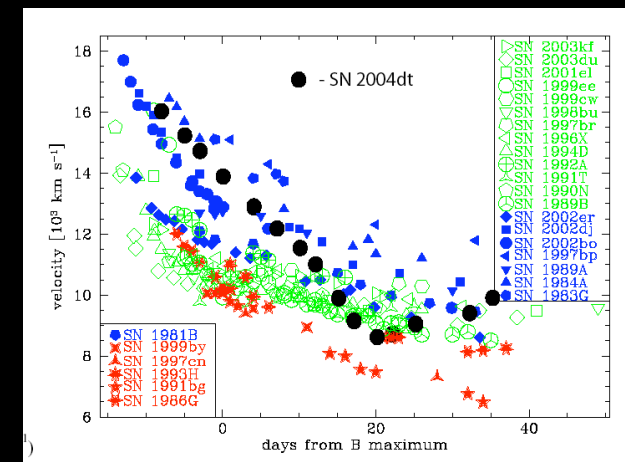
## Luminosity Indicators



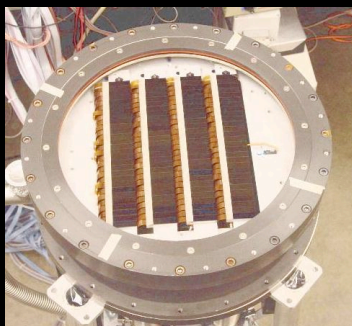
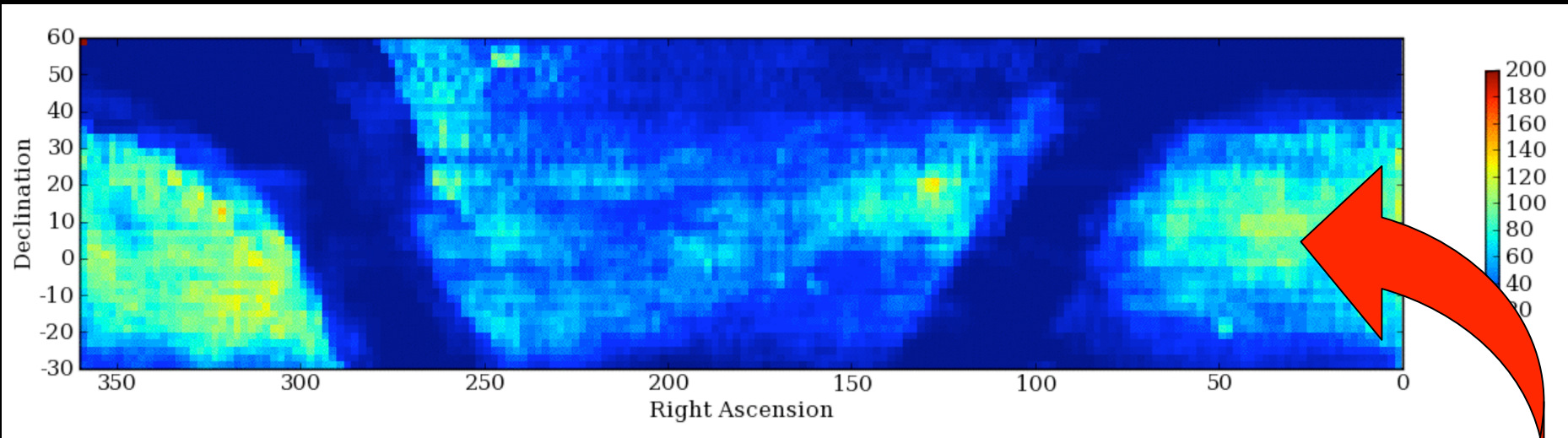
## Abundance Tomography



## Kinematics

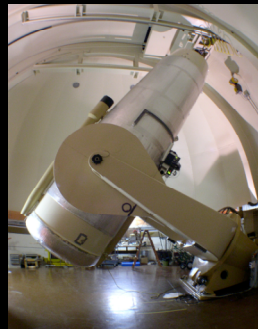


# Nearby Supernova Factory Search



QUEST Camera

+



Oschin Schmidt

+



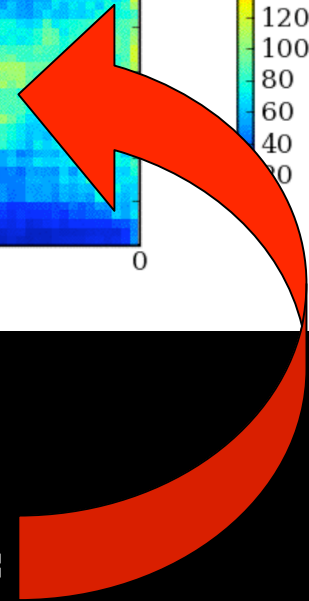
HPWREN

+

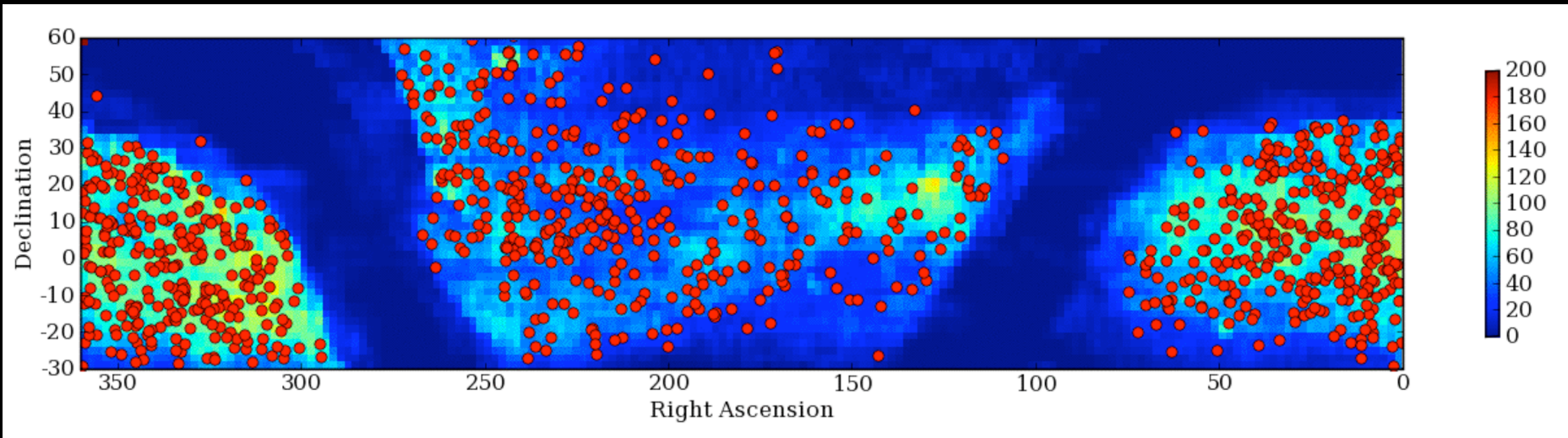


NERSC

=



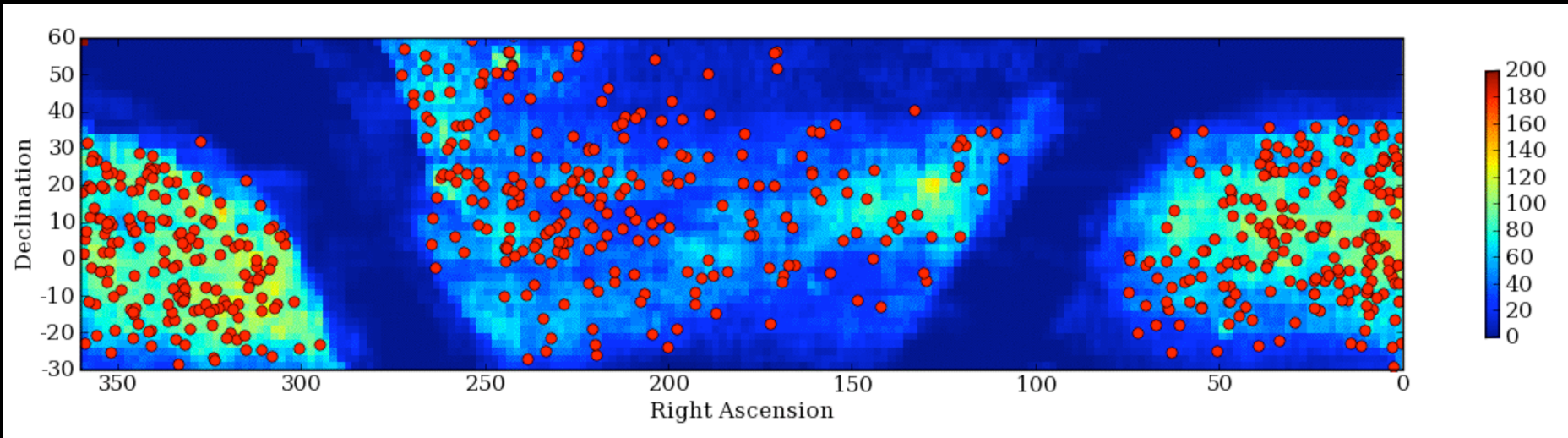
# SNe over $2\pi$ of the Sky



More than 1000 SNe in  $\sim 28$  months of searching  
during 2005-2008

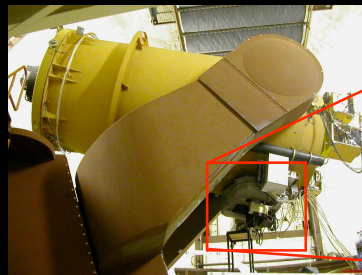


# SNe over $2\pi$ of the Sky



624 SNe w/ spectroscopic types - mostly from the  
**SuperNova Integral Field Spectrograph**

UH88



SNIFS



396 SNfactory & 50 IAUC SNe Ia

# We faced fire and brimstone to get the data!



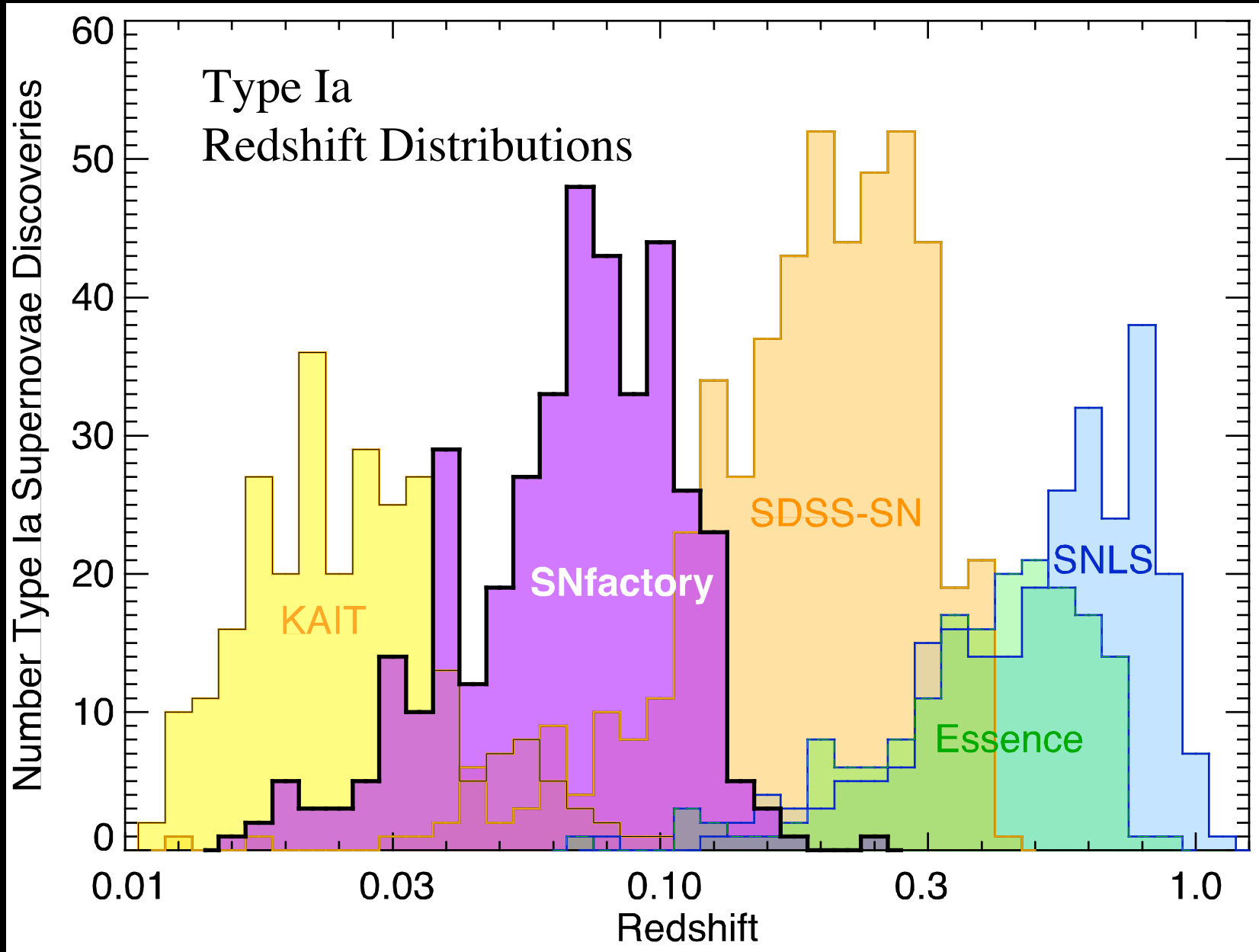
Fire closes  
Palomar Observatory



Mag 6.7 earthquake hits Mauna Kea

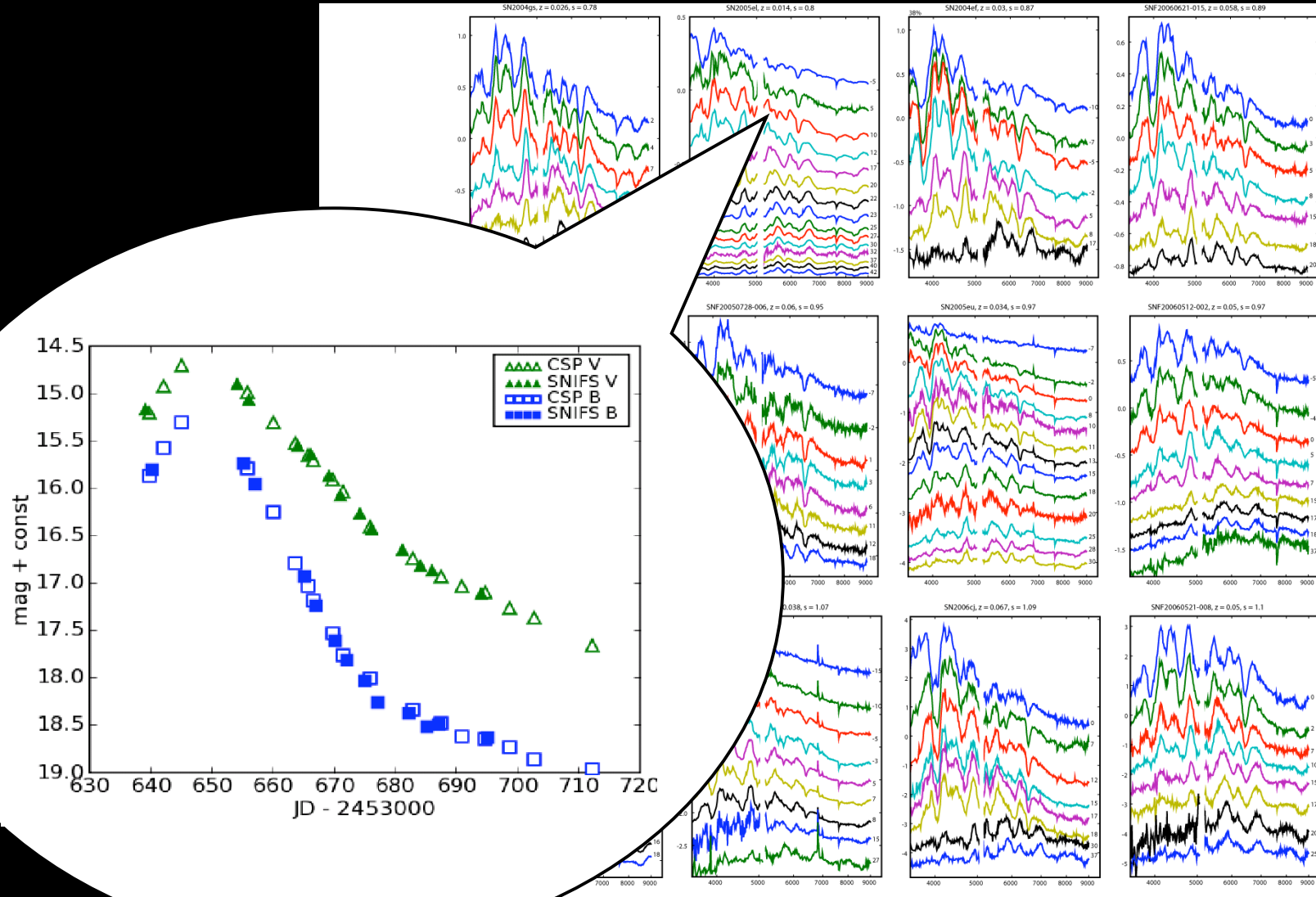


Fire sweeps antenna site at North Peak



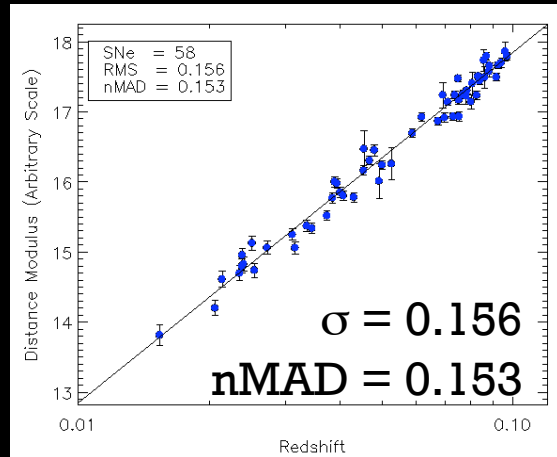


# Library of Spectral Time Series

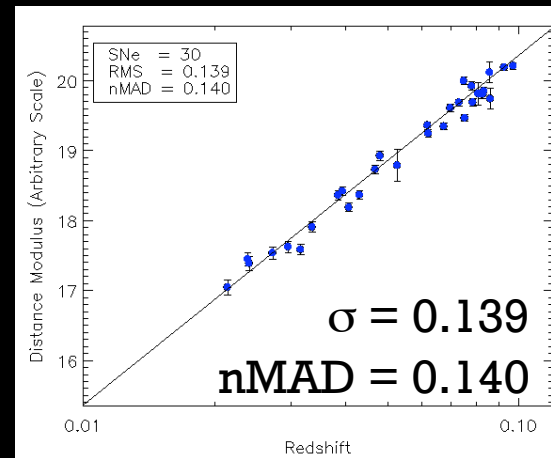


# Now Several Paths to SN Hubble Diagram

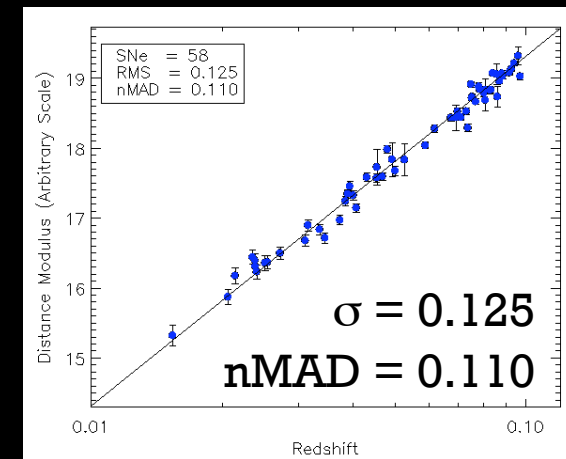
## Stretch-Color



## $\mathcal{R}(\text{SiSS})$ -Color

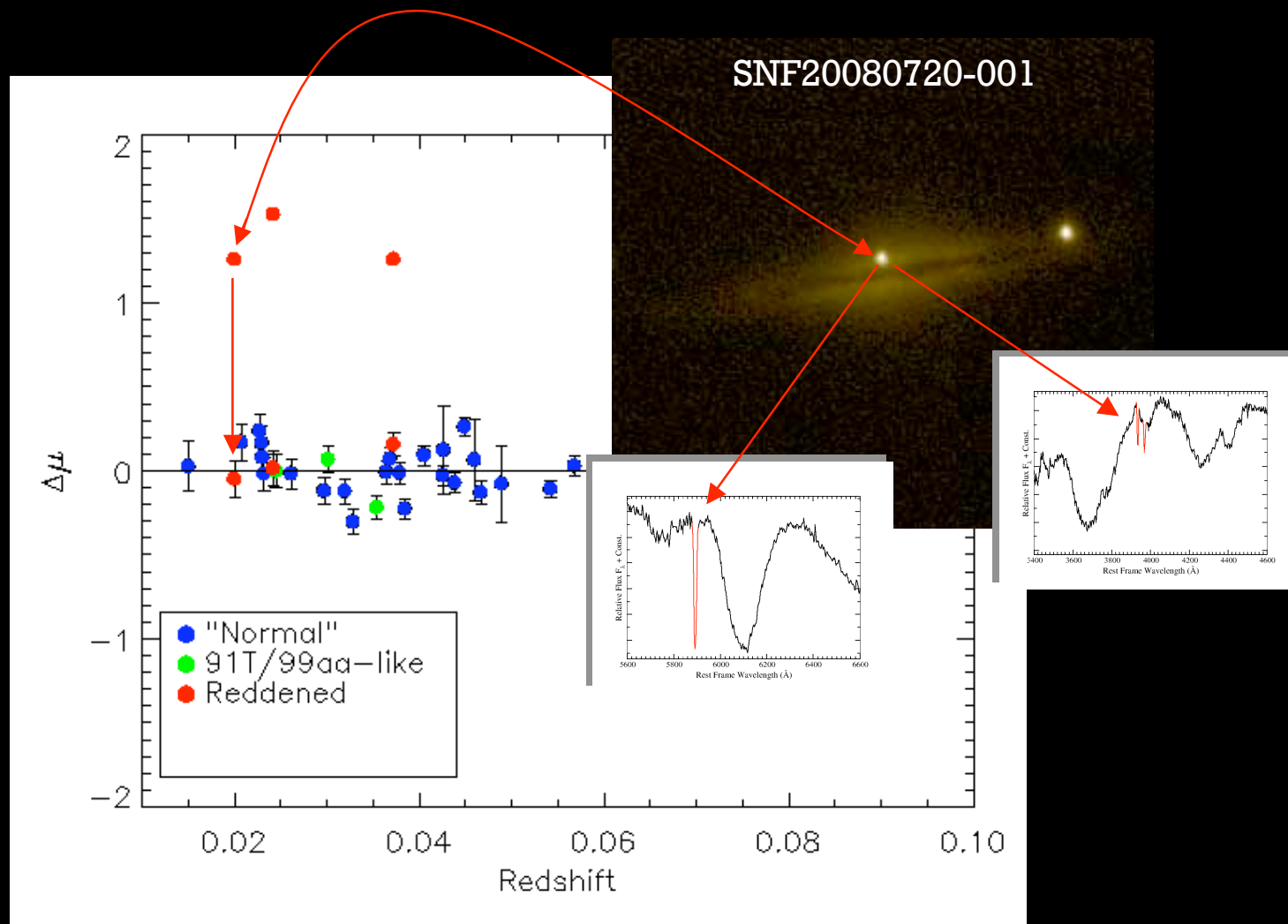


## $\mathcal{R}(642/443)$



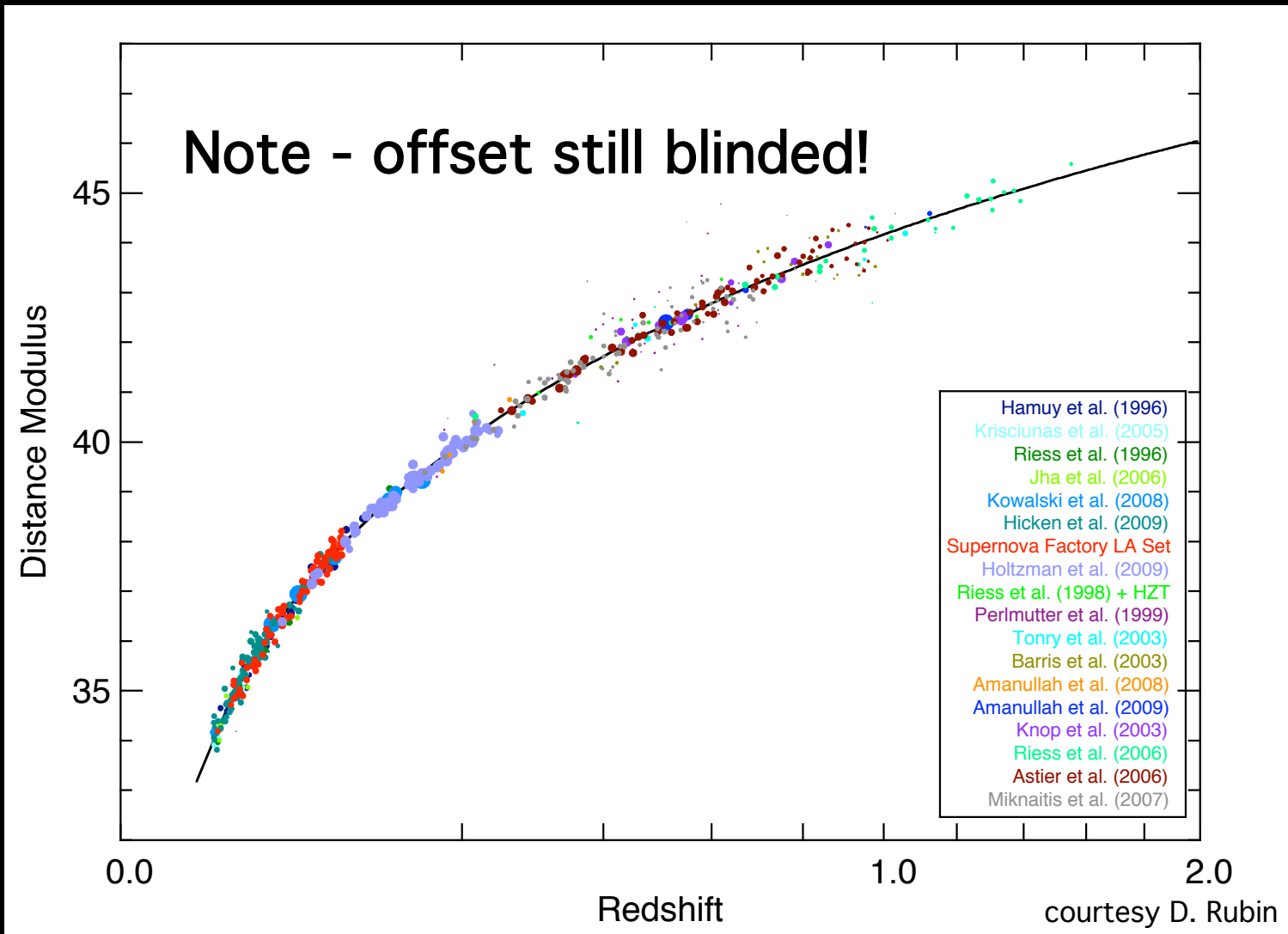
Bailey et al, 2009

# $R_{642/443}$ Handles Difficult Cases

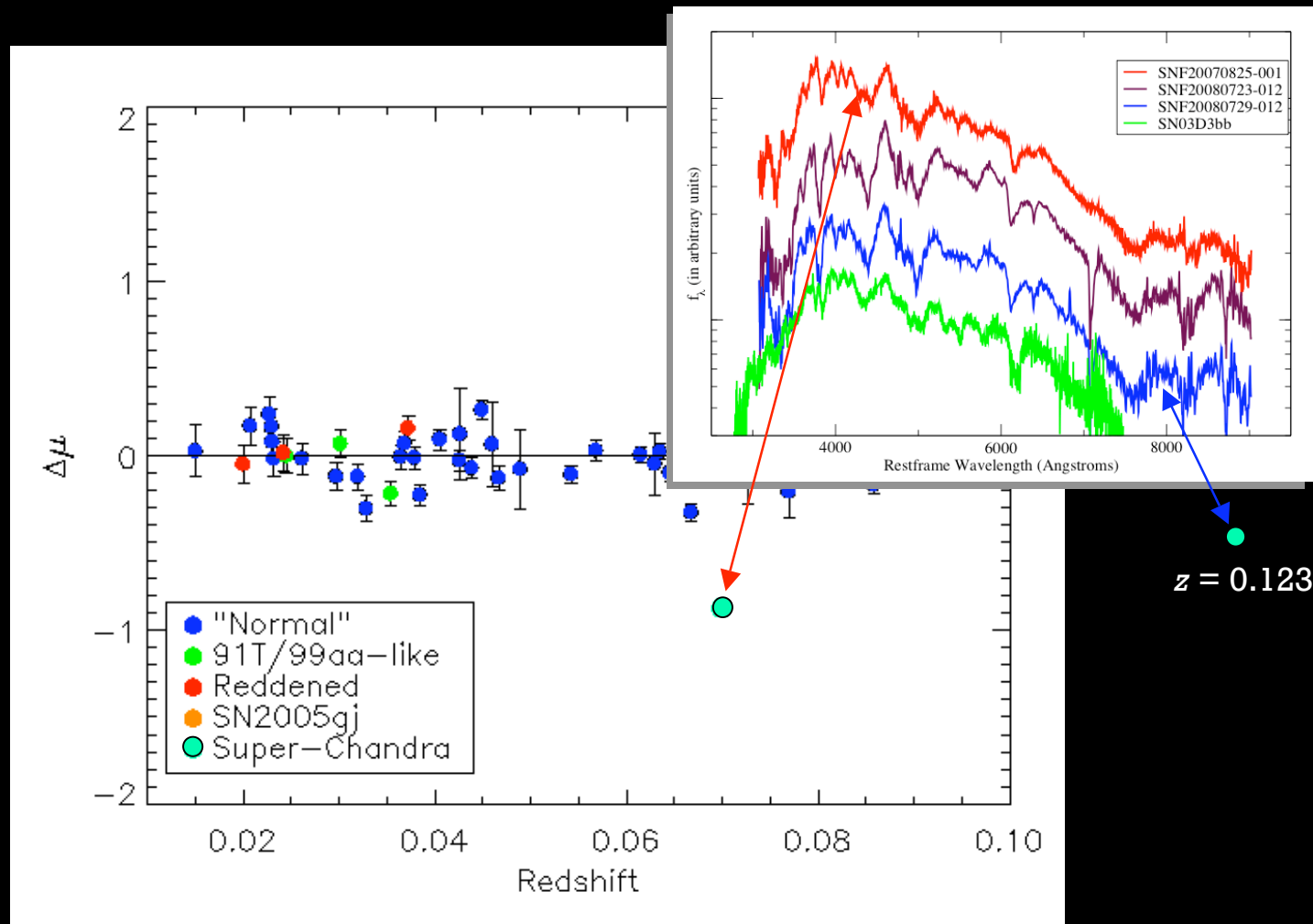




# Stretch-Color Hubble Diagram

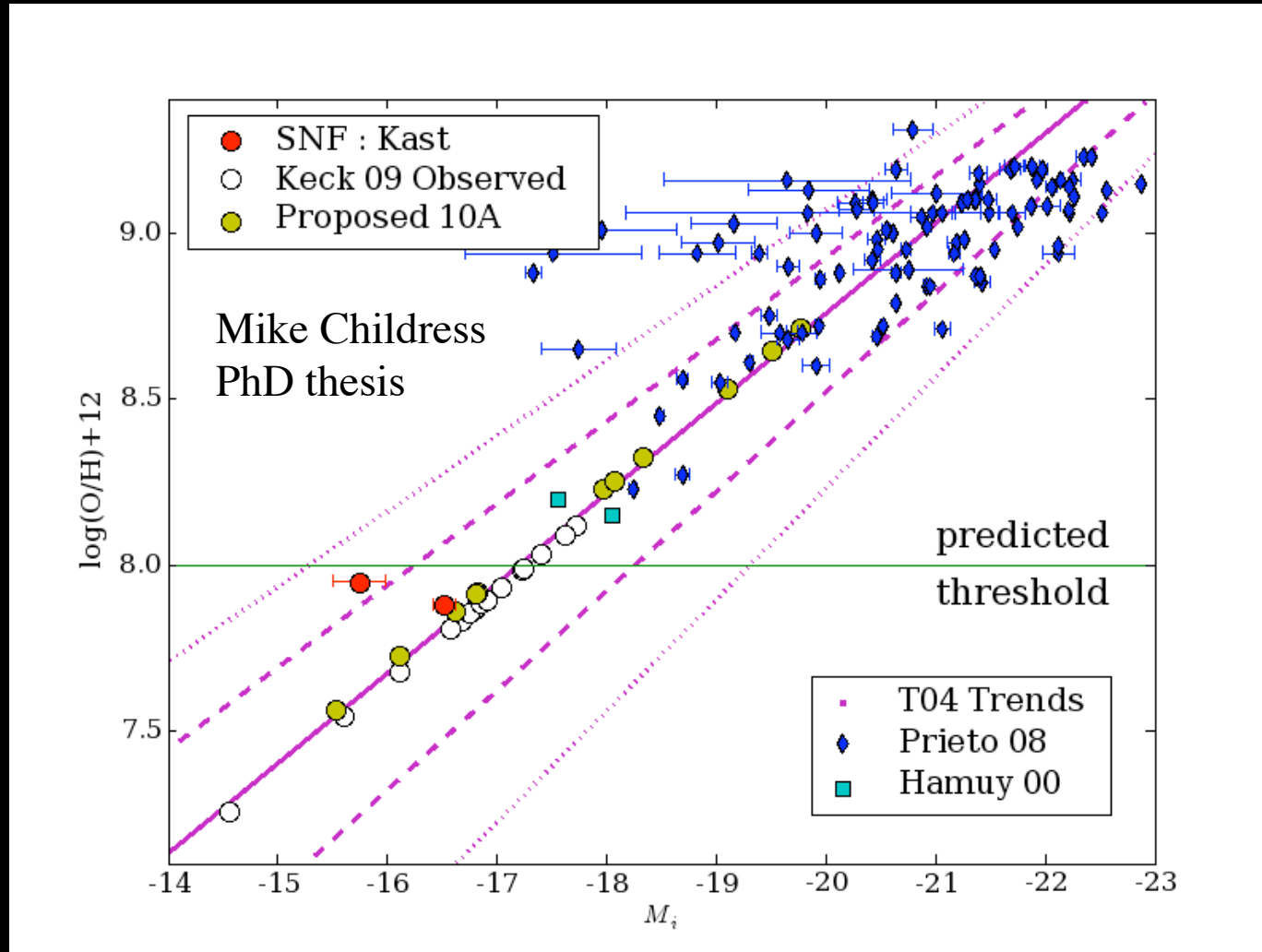


# Have We Seen the Double-Degenerates?



# Environmental Clues and Constraints

Galaxy Metallicity  $\uparrow$



Galaxy Luminosity  $\longrightarrow$



A/D converter module

Photodiode preamp

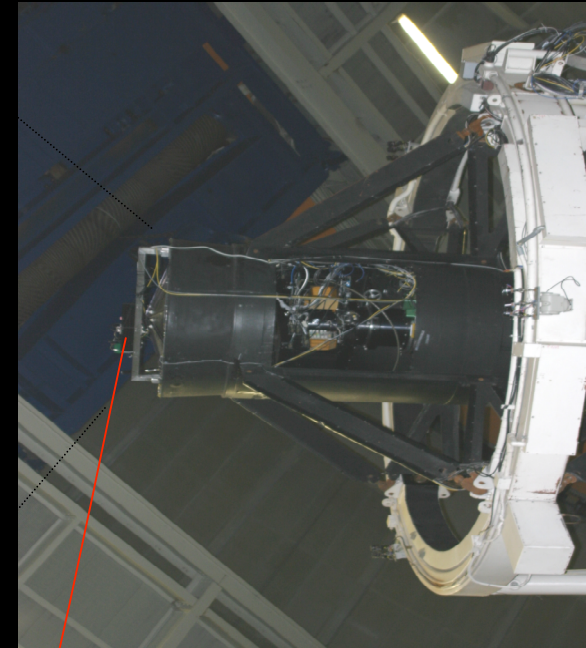
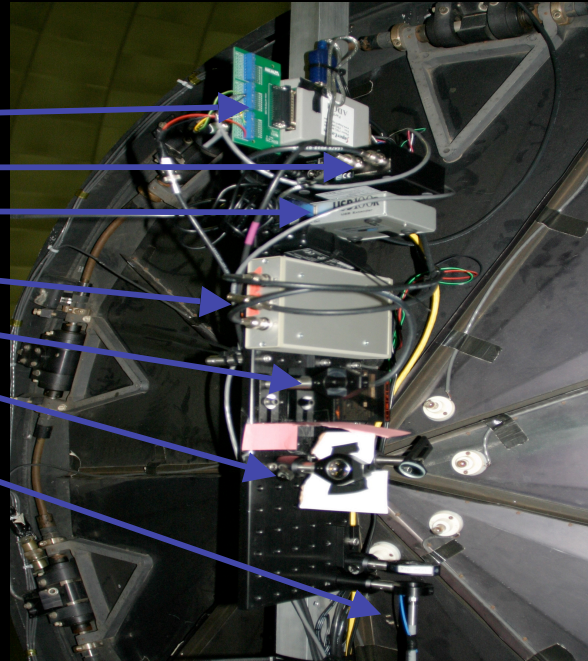
USB extender

Integrator electronics

Calibrated photodiode

Beam launch optics

Optical Fiber



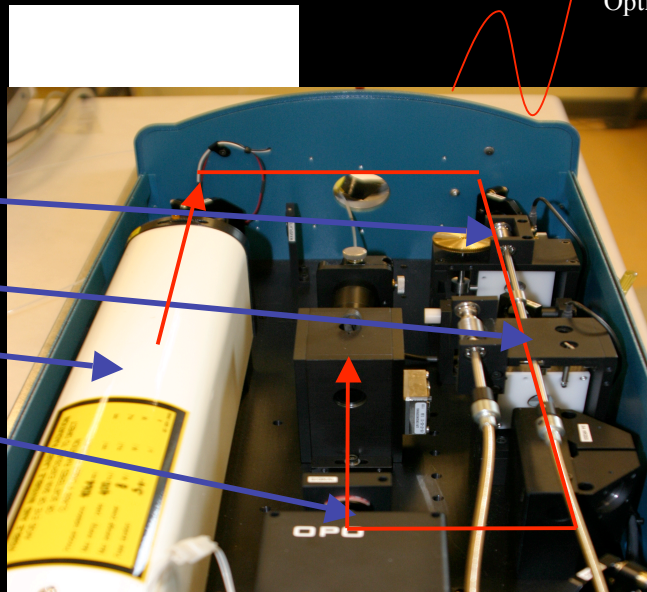
Optical fiber

Second harmonic (532 nm) generator

Mixer (to 355 nm)

1.06 micron Nd:YAG pulsed laser

Tunable downconverter



Technique:

- Illuminate flat field screen with monochromatic laser light from a tunable laser, and take a spectral image at each  $\lambda$ .
- Deliver a fixed “dose” of light to the system, as measured by a NIST calibrated photodiode.
- Measuring the flatfield flux relative to that seen by the photodiode allows the measurement of the total system throughput, in situ, relative to the well-characterized diode.

Courtesy C. Stubbs

# Work for the Coming Year

- SNIFS
  - Complete final references for 2005-2008 SNe
  - Follow-up of SNe Ia targeted by HST/STIS
  - Supplement current sample with other interesting SNe
  - Development of absolute color calibration system
- Cosmology analysis
  - Finalize flux calibration
  - Complete host galaxy redshifts
  - Finalize host subtraction pipeline
- Spectral analysis
  - Develop new/improved spectral luminosity indicators
  - Abundance tomography
  - Spectral temporal-luminosity model,  $F(\lambda, t, s)$



# Summary

- Several fundamentally different standardization methods have been demonstrated to work well using spectrophotometric data.
- New methods using spectrophotometry are showing significant improvement over imaging based methods.
- Type Ia population dominated by SNe standardizable to 6%.
- Highly-extinct SNe Ia can be standardized at optical wavelengths
  - more examples needed to break intrinsic color - dust degeneracy
- Oddball SNe Ia and those in low-luminosity hosts help probe progenitor properties.
- Deeper understanding of SNe Ia for cosmology demands spectrophotometry

